

Statistical Thermodynamics

I. Introduction: Temperature of an Noble Gas

Further reading: Quanta, Matter and Change, Peter Atkins

1. State of a system

Thermodynamic system: part of Universe under consideration.
→ real/imaginary boundary separates from rest, i.e. environment

- * isolated: ϕ heat/matter/work exchange
- * closed: ϕ matter exchange, $q + w$
 - adiabatic: ϕ q exchange $\neq$ - rigid: ϕ w exchange.
- * open: free exchange with environment

State: System in eq. under certain conditions = state
→ System properties have specific value, e.g. composition, p, V, T
* function of state & independent of path

Ideal gas equation:

$$pV = nRT$$

→ Avogadro's law

For an ideal gas: $T \propto \sum K_E$, i.e.

$$\frac{1}{2} M c^2 = \frac{3}{2} RT$$

Proof: $\bullet V_x^2 = \frac{1}{3} c^2$ • mean free path: ℓ • time $\tau = \frac{\ell}{V_x}$
→ n° collisions = $\frac{1}{2} \ell A p$ (in time τ)

$$\begin{aligned} p &= \frac{F}{A} \quad F = \frac{\Delta m v}{\Delta t} \Rightarrow p = \frac{A \ell p}{2} \cdot \frac{2 m v_x}{\tau} \cdot \frac{1}{A} \Rightarrow p = \frac{A \ell p^2 v_x^2 m}{2 \ell A} \\ \Leftrightarrow p &= \frac{1}{3} \rho m c^2 \Leftrightarrow pV = \frac{1}{3} N m c^2 \\ \Leftrightarrow pV &= \frac{1}{3} N M c^2 \\ \Leftrightarrow RT &= \frac{1}{3} M c^2 \\ \Leftrightarrow \frac{1}{2} M c^2 &= \frac{3}{2} RT \end{aligned}$$

Boyle's law: $p \propto \frac{1}{V}$

Charles' law: $V \propto T$ Avogadro's law: $n \propto T$