Gas molecules attract one another and stick together, as the IMFs become significant, and there are fewer particles bouncing around and creating pressure, making the pressure of a nonideal gas smaller than the pressure predicted by the ideal gas equation

Stronger IMFs will lead to more deviations

→ H<sub>2</sub>O, with hydrogen bonding, is more likely to deviate from ideal behavior than CH<sub>4</sub>, with only London dispersion forces

When gases have similar IMFs, gases with more electrons are more polarizable and more likely to deviate from ideal behavior

→ For noble gases with only LDFs, argon is more likely to deviate than helium, but less likely to deviate than xenon

## DENSITY

Measured in the same way of a liquid or solid, D = m/v

D = density, m = mass of gas (grams), V = volume occupied by gas (liters)

Determine density of a gas sample by combining density equation density Gas Law

- → Substitute V = m/D and rearrange until D isolated

Therefore, 
$$PV = \frac{m}{M} RT$$
  
and  $\rho = \frac{m}{V} = \frac{PM}{RT}$ 

→

## SPECTROSCOPY AND THE EM SPECTRUM

Differences in absorption or emission of photons in different regions are related to different types of molecular motion or electronic transition:

- → Microwave radiation transitions in molecular rotational levels
- → Infrared radiation transitions in molecular vibrational levels
- → Ultraviolet and visible radiation transitions in electronic energy levels

## PHOTOELECTRIC EFFECT

When an atom or molecule absorbs/emits a photon, the energy of the species is increased/decreased by an amount equal to the photon's energy Equation relates wavelength of electromagnetic wave to its frequency and the speed of

light