

* Ionization Energy :-

→ energy required to sending e^- to infinity.

$$\Delta E = I.E = +13.6 Z^2 \text{ eV}$$

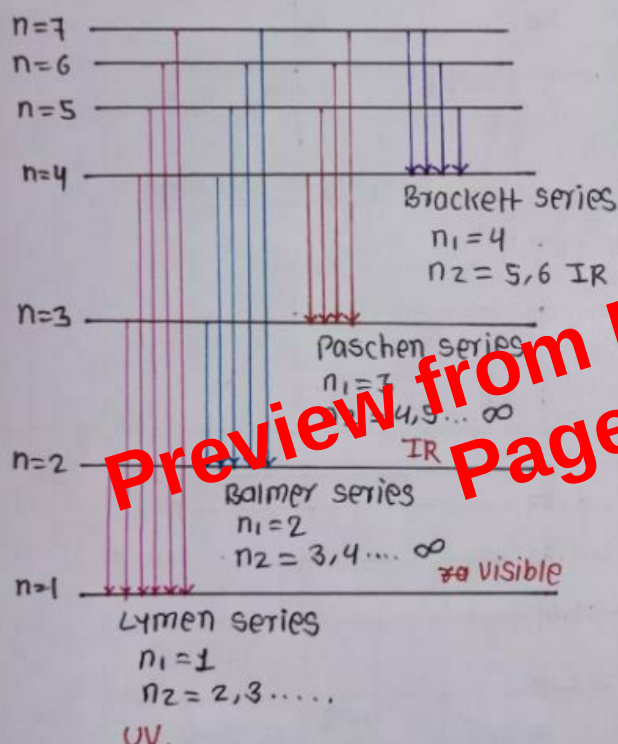
$$\Delta E = I.E = +13.6 Z^2 \times 96 \text{ KJ/mol}$$

* Binding energy :-

→ Energy required to ~~some~~ e^- remove e^- from given energy shell.

$$BE = | \text{energy of } n\text{th shell} | = 13.6 Z^2 / n^2 \text{ eV}$$

* Hydrogen Spectrum *



* Rydberg's Formula *

$$\vec{v} = \frac{1}{\lambda} = RZ^2 \left[\frac{1}{n_1^2} - \frac{1}{n_2^2} \right]$$

$$R = 109677 \text{ cm}^{-1} \approx 1.0967 \times 10^5 = 10^5 \text{ cm}^{-1}$$

* Lyman series $\rightarrow (\alpha \text{ or } H_\alpha \text{ line})$

$$\lambda_\alpha = \frac{4}{3RZ^2} \text{ — First line of Lyman}$$

→ Second line (β or H_β line)

$$\lambda_\beta = \frac{9}{8RZ^2}$$

→ Line of shortest wavelength

$$\lambda_{\min} = \frac{1}{RZ^2}$$

→ Line of longest wavelength

$$\lambda_\alpha = \frac{4}{3RZ^2}$$

• calculation of no. of spectral line:

$$\frac{(n_2 - n_1)(n_2 - n_1 + 1)}{2}$$

$$n_2 - n_1$$

n_2 — final state of that series

Specific series.

Drawback of Bohr's model

- Applicable for single e^- species
- not explain Stark effect & Zeeman effect.

• de-Broglie concept :-

$$\lambda \propto \frac{1}{p}$$

$$\lambda \propto \frac{h}{p} = \frac{h}{mv} = \frac{h}{\sqrt{2m(K.E.)}}$$

$$\lambda_e = \sqrt{\frac{150}{v}} \text{ \AA} \text{ or } \frac{12.27}{\sqrt{v}} \text{ \AA}$$



particle accelerated from rest.